

SAS/ETS[®] 14.1 User's Guide

The SPECTRA Procedure

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Chapter 33

The SPECTRA Procedure

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Overview: SPECTRA Procedure

The SPECTRA procedure performs spectral and cross-spectral analysis of time series. You can use spectral analysis techniques to look for periodicities or cyclical patterns in data.

The SPECTRA procedure produces estimates of the spectral and cross-spectral densities of a multivariate time series. Estimates of the spectral and cross-spectral densities of a multivariate time series are produced using a finite Fourier transform to obtain periodograms and cross-periodograms. The periodogram ordinates are smoothed by a moving average to produce estimated spectral and cross-spectral densities. PROC SPECTRA can also test whether or not the data are white noise.

PROC SPECTRA uses the finite Fourier transform to decompose data series into a sum of sine and cosine waves of different amplitudes and wavelengths. The finite Fourier transform decomposition of the series x_t is

$$x_t = \frac{a_0}{2} + \sum_{k=1}^{m-1} f_k (a_k \cos \omega_k t + b_k \sin \omega_k t)$$

$$f_k = \begin{cases} 1/2 & \text{if } n \text{ is even and } k = m - 1 \\ 1 & \text{otherwise} \end{cases}$$

where

t	is the time subscript, $t = 0, 1, 2, \dots, n - 1$
x_t	are the equally spaced time series data
n	is the number of observations in the time series
m	is the number of frequencies in the Fourier decomposition: $m = \frac{n+2}{2}$ if n is even, $m = \frac{n+1}{2}$ if n is odd
k	is the frequency subscript, $k = 0, 1, 2, \dots, m - 1$
a_0	is the mean term: $a_0 = 2\bar{x}$
a_k	are the cosine coefficients
b_k	are the sine coefficients
ω_k	are the Fourier frequencies: $\omega_k = \frac{2\pi k}{n}$

Functions of the Fourier coefficients a_k and b_k can be plotted against frequency or against wave length to form *periodograms*. The amplitude periodogram J_k is defined as follows:

$$J_k = \frac{n}{2} (a_k^2 + b_k^2)$$

Several definitions of the term periodogram are used in the spectral analysis literature. The following discussion refers to the J_k sequence as the periodogram.

The periodogram can be interpreted as the contribution of the k th harmonic ω_k to the total sum of squares (in an analysis of variance sense) in the decomposition of the process into two-degree-of-freedom components for each of the m frequencies. When n is even, $\sin(\omega_{\frac{n}{2}})$ is zero, and thus the last periodogram value is a one-degree-of-freedom component.

The periodogram is a volatile and inconsistent estimator of the spectrum. The spectral density estimate is produced by smoothing the periodogram. Smoothing reduces the variance of the estimator but introduces a bias. The weight function used for the smoothing process, $W()$, often called the kernel or spectral window, is specified with the WEIGHTS statement. It is related to another weight function, $w()$, the lag window, that is used in other methods to taper the correlogram rather than to smooth the periodogram. Many specific weighting functions have been suggested in the literature (Fuller 1976; Jenkins and Watts 1968; Priestley 1981). Table 33.3 later in this chapter gives the relevant formulas when the WEIGHTS statement is used.

Letting i represent the imaginary unit $\sqrt{-1}$, the cross-periodogram is defined as follows:

$$J_k^{xy} = \frac{n}{2} (a_k^x a_k^y + b_k^x b_k^y) + i \frac{n}{2} (a_k^x b_k^y - b_k^x a_k^y)$$

The cross-spectral density estimate is produced by smoothing the cross-periodogram in the same way as the periodograms are smoothed using the spectral window specified by the WEIGHTS statement.

The SPECTRA procedure creates an output SAS data set whose variables contain values of the periodograms, cross-periodograms, estimates of spectral densities, and estimates of cross-spectral densities. The form of the output data set is described in the section “OUT= Data Set” on page 2339.

Getting Started: SPECTRA Procedure

To use the SPECTRA procedure, specify the input and output data sets and options for the analysis you want in the PROC SPECTRA statement, and list the variables to analyze in the VAR statement. The procedure produces no printed output unless the WHITEST option is specified in the PROC SPECTRA statement. The periodogram, spectral density, and other results are written to the OUT= data set, depending on the options used.

For example, to compute the Fourier transform of a variable X in a data set A, use the following statements:

```
proc spectra data=a out=b coef;
  var x;
run;
```

This PROC SPECTRA step writes the Fourier coefficients a_k and b_k to the variables COS_01 and SIN_01 in the output data set B.

When a WEIGHTS statement is specified, the periodogram is smoothed by a weighted moving average to produce an estimate of the spectral density of the series. The following statements write a spectral density estimate for X to the variable S_01 in the output data set B.

```
proc spectra data=a out=b s;
  var x;
  weights 1 2 3 4 3 2 1;
run;
```

When the VAR statement specifies more than one variable, you can perform cross-spectral analysis by specifying the CROSS option in the PROC SPECTRA statement. The CROSS option by itself produces the cross-periodograms for all two-way combinations of the variables listed in the VAR statement. For example, the following statements write the real and imaginary parts of the cross-periodogram of X and Y to the variables RP_01_02 and IP_01_02 in the output data set B.

```
proc spectra data=a out=b cross;
  var x y;
run;
```

To produce cross-spectral density estimates, specify both the CROSS option and the S option. The cross-periodogram is smoothed using the weights specified by the WEIGHTS statement in the same way as the spectral density. The squared coherency and phase estimates of the cross-spectrum are computed when the K and PH options are used.

The following example computes cross-spectral density estimates for the variables X and Y.

```
proc spectra data=a out=b cross s;
  var x y;
  weights 1 2 3 4 3 2 1;
run;
```

The real part and imaginary part of the cross-spectral density estimates are written to the variables CS_01_02 and QS_01_02, respectively.

Syntax: SPECTRA Procedure

The following statements are used with the SPECTRA procedure:

```
PROC SPECTRA options ;
BY variables ;
VAR variables ;
WEIGHTS < weights> < kernel> ;
```

Functional Summary

Table 33.1 summarizes the statements and options that control the SPECTRA procedure.

Table 33.1 SPECTRA Functional Summary

Description	Statement	Option
Statements		
specify BY-group processing	BY	
specify the variables to be analyzed	VAR	
specify weights for spectral density estimates	WEIGHTS	
Data Set Options		
specify the input data set	PROC SPECTRA	DATA=
specify the output data set	PROC SPECTRA	OUT=
Output Control Options		
output the amplitudes of the cross-spectrum	PROC SPECTRA	A
output the Fourier coefficients	PROC SPECTRA	COEF
output the periodogram	PROC SPECTRA	P
output the spectral density estimates	PROC SPECTRA	S
output cross-spectral analysis results	PROC SPECTRA	CROSS
output squared coherency of the cross-spectrum	PROC SPECTRA	K
output the phase of the cross-spectrum	PROC SPECTRA	PH

Table 33.1 *continued*

Description	Statement	Option
Smoothing Options		
specify the Bartlett kernel	WEIGHTS	BART
specify the Parzen kernel	WEIGHTS	PARZEN
specify the quadratic spectral kernel	WEIGHTS	QS
specify the Tukey-Hanning kernel	WEIGHTS	TUKEY
specify the truncated kernel	WEIGHTS	TRUNCAT
Other Options		
subtract the series mean	PROC SPECTRA	ADJMEAN
specify an alternate quadrature spectrum estimate	PROC SPECTRA	ALTW
request tests for white noise	PROC SPECTRA	WHITESTEST

PROC SPECTRA Statement

PROC SPECTRA *options* ;

The following options can be used in the PROC SPECTRA statement:

A

outputs the amplitude variables (A_{nn_mm}) of the cross-spectrum.

ADJMEAN

CENTER

subtracts the series mean before performing the Fourier decomposition. This sets the first periodogram ordinate to 0 rather than $2n$ times the squared mean. This option is commonly used when the periodograms are to be plotted to prevent a large first periodogram ordinate from distorting the scale of the plot.

ALTW

specifies that the quadrature spectrum estimate is computed at the boundaries in the same way as the spectral density estimate and the cospectrum estimate are computed.

COEF

outputs the Fourier cosine and sine coefficients of each series.

CROSS

is used with the P and S options to output cross-periodograms and cross-spectral densities when more than one variable is listed in the VAR statement.

DATA=SAS-data-set

names the SAS data set that contains the input data. If the DATA= option is omitted, the most recently created SAS data set is used.

K

outputs the squared coherency variables (K_{nn_mm}) of the cross-spectrum. The K_{nn_mm} variables are identically 1 unless weights are given in the WEIGHTS statement and the S option is specified.

OUT=SAS-data-set

names the output data set created by PROC SPECTRA to store the results. If the OUT= option is omitted, the output data set is named by using the DATA n convention.

P

outputs the periodogram variables. The variables are named P_{nn} , where nn is an index of the original variable with which the periodogram variable is associated. When both the P and CROSS options are specified, the cross-periodogram variables RP_{nn_mm} and IP_{nn_mm} are also output.

PH

outputs the phase variables (PH_{nn_mm}) of the cross-spectrum.

S

outputs the spectral density estimates. The variables are named S_{nn} , where nn is an index of the original variable with which the estimate variable is associated. When both the S and CROSS options are specified, the cross-spectral variables CS_{nn_mm} and QS_{nn_mm} are also output.

WHITETEST

prints two tests of the hypothesis that the data are white noise. See the section “[White Noise Test](#)” on page 2339 for details.

Note that the CROSS, A, K, and PH options are meaningful only if more than one variable is listed in the VAR statement.

BY Statement

BY *variables* ;

A BY statement can be used with PROC SPECTRA to obtain separate analyses for groups of observations defined by the BY variables.

VAR Statement

VAR *variables* ;

The VAR statement specifies one or more numeric variables that contain the time series to analyze. The order of the variables in the VAR statement list determines the index, nn , used to name the output variables. The VAR statement is required.

WEIGHTS Statement

WEIGHTS *weight-constants / kernel-specification* ;

The WEIGHTS statement specifies the relative weights used in the moving average applied to the periodogram ordinates to form the spectral density estimates. A WEIGHTS statement must be used to produce smoothed spectral density estimates. You can specify the relative weights in two ways: you can specify them explicitly as explained in the section “[Using Weight Constants Specification](#)” on page 2335, or you can specify them implicitly by using the kernel specification as explained in the section “[Using Kernel Specifications](#)” on page 2335. If the WEIGHTS statement is not used, only the periodogram is produced.

Using Weight Constants Specification

Any number of weighting constants can be specified. The constants should be positive and symmetric about the middle weight. The middle constant (or the constant to the right of the middle if an even number of weight constants are specified) is the relative weight of the current periodogram ordinate. The constant immediately following the middle one is the relative weight of the next periodogram ordinate, and so on. The actual weights used in the smoothing process are the weights specified in the WEIGHTS statement scaled so that they sum to $\frac{1}{4\pi}$.

The moving average reflects at each end of the periodogram. The first periodogram ordinate is not used; the second periodogram ordinate is used in its place.

For example, a simple triangular weighting can be specified using the following WEIGHTS statement:

```
weights 1 2 3 2 1;
```

Using Kernel Specifications

You can specify five different kernels in the WEIGHTS statement. The syntax for the statement is

WEIGHTS [PARZEN][BART][TUKEY][TRUNCAT][QS] [c e] ;

where $c \geq 0$ and $e \geq 0$ are used to compute the bandwidth parameter as

$$l(q) = cq^e$$

and q is the number of periodogram ordinates + 1:

$$q = \text{floor}(n/2) + 1$$

To specify the bandwidth explicitly, set $c =$ to the desired bandwidth and $e = 0$.

For example, a Parzen kernel can be specified using the following WEIGHTS statement:

```
weights parzen 0.5 0;
```

For details, see the section “[Kernels](#)” on page 2336.

Details: SPECTRA Procedure

Input Data

Observations in the data set analyzed by the SPECTRA procedure should form ordered, equally spaced time series. No more than 99 variables can be included in the analysis.

Data are often detrended before analysis by the SPECTRA procedure. This can be done by using the residuals output by a SAS regression procedure. Optionally, the data can be centered using the ADJMEAN option in the PROC SPECTRA statement, since the zero periodogram ordinate corresponding to the mean is of little interest from the point of view of spectral analysis.

Missing Values

Missing values are excluded from the analysis by the SPECTRA procedure. If the SPECTRA procedure encounters missing values for any variable listed in the VAR statement, the procedure determines the longest contiguous span of data that has no missing values for the variables listed in the VAR statement and uses that span for the analysis.

Computational Method

If the number of observations n factors into prime integers that are less than or equal to 23, and the product of the square-free factors of n is less than 210, then PROC SPECTRA uses the fast Fourier transform developed by Cooley and Tukey and implemented by Singleton (1969). If n cannot be factored in this way, then PROC SPECTRA uses a Chirp-Z algorithm similar to that proposed by Monro and Branch (1977). To reduce memory requirements, when n is small, the Fourier coefficients are computed directly using the defining formulas.

Kernels

Kernels are used to smooth the periodogram by using a weighted moving average of nearby points. A smoothed periodogram is defined by the following equation.

$$\hat{J}_i(l(q)) = \sum_{\tau=-l(q)}^{l(q)} w\left(\frac{\tau}{l(q)}\right) \tilde{J}_{i+\tau}$$

where $w(x)$ is the kernel or weight function. At the endpoints, the moving average is computed cyclically; that is,

$$\tilde{J}_{i+\tau} = \begin{cases} J_{i+\tau} & 0 \leq i + \tau \leq q \\ J_{-(i+\tau)} & i + \tau < 0 \\ J_{q-(i+\tau)} & i + \tau > q \end{cases}$$

The SPECTRA procedure supports the following kernels. They are listed with their default bandwidth functions.

Bartlett: KERNEL BART

$$\begin{aligned} w(x) &= \begin{cases} 1 - |x| & |x| \leq 1 \\ 0 & \text{otherwise} \end{cases} \\ l(q) &= \frac{1}{2}q^{1/3} \end{aligned}$$

Parzen: KERNEL PARZEN

$$\begin{aligned} w(x) &= \begin{cases} 1 - 6|x|^2 + 6|x|^3 & 0 \leq |x| \leq \frac{1}{2} \\ 2(1 - |x|)^3 & \frac{1}{2} \leq |x| \leq 1 \\ 0 & \text{otherwise} \end{cases} \\ l(q) &= q^{1/5} \end{aligned}$$

Quadratic spectral: KERNEL QS

$$\begin{aligned} w(x) &= \frac{25}{12\pi^2 x^2} \left(\frac{\sin(6\pi x/5)}{6\pi x/5} - \cos(6\pi x/5) \right) \\ l(q) &= \frac{1}{2}q^{1/5} \end{aligned}$$

Tukey-Hanning: KERNEL TUKEY

$$\begin{aligned} w(x) &= \begin{cases} (1 + \cos(\pi x))/2 & |x| \leq 1 \\ 0 & \text{otherwise} \end{cases} \\ l(q) &= \frac{2}{3}q^{1/5} \end{aligned}$$

Truncated: KERNEL TRUNCAT

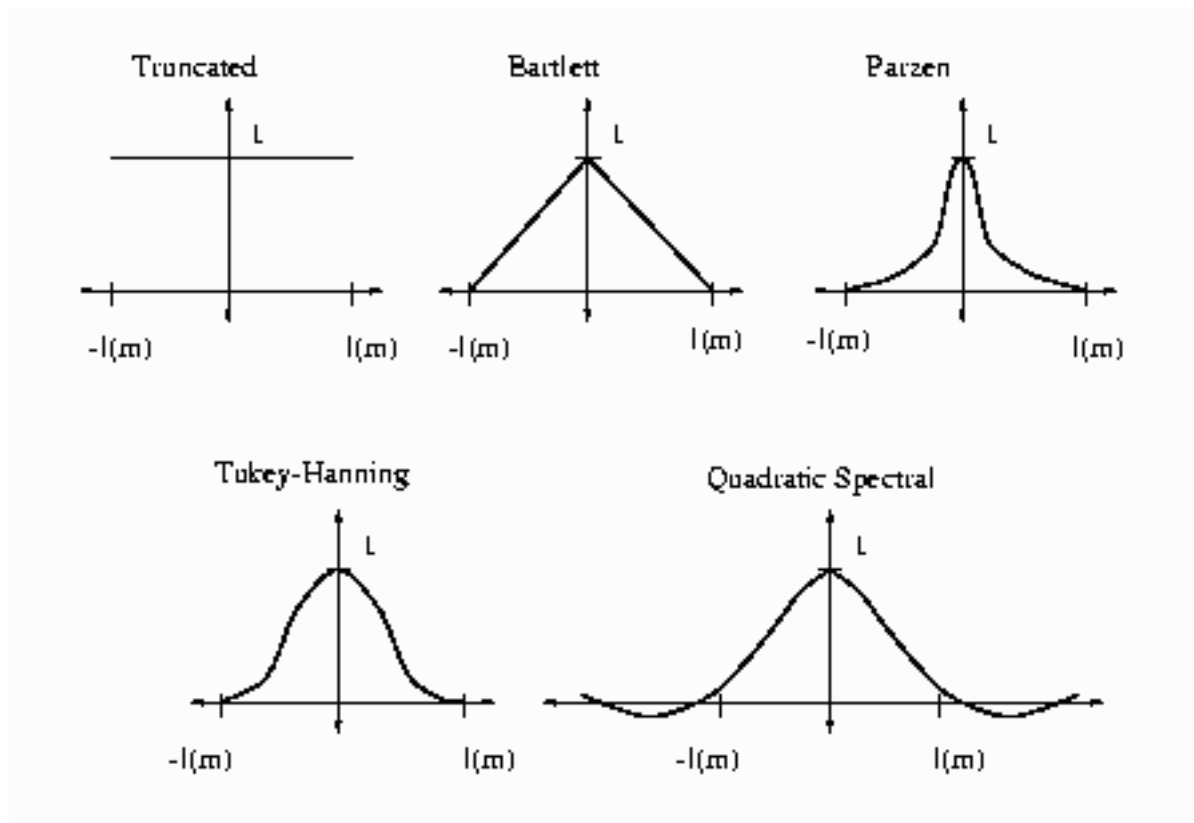
$$\begin{aligned} w(x) &= \begin{cases} 1 & |x| \leq 1 \\ 0 & \text{otherwise} \end{cases} \\ l(q) &= \frac{1}{4}q^{1/5} \end{aligned}$$

A summary of the default values of the bandwidth parameters, c and e , associated with the kernel smoothers in PROC SPECTRA are listed below in Table 33.2:

Table 33.2 Bandwidth Parameters

Kernel	c	e
Bartlett	$1/2$	$1/3$
Parzen	1	$1/5$
quadratic	$1/2$	$1/5$
Tukey-Hanning	$2/3$	$1/5$
truncated	$1/4$	$1/5$

Figure 33.1 Kernels for Smoothing



See Andrews (1991) for details about the properties of these kernels.

White Noise Test

PROC SPECTRA prints two test statistics for white noise when the WHITETEST option is specified: Fisher's Kappa (Davis 1941; Fuller 1976) and Bartlett's Kolmogorov-Smirnov statistic (Bartlett 1966; Fuller 1976; Durbin 1967).

If the time series is a sequence of independent random variables with mean 0 and variance σ^2 , then the periodogram, J_k , will have the same expected value for all k . For a time series with nonzero autocorrelation, each ordinate of the periodogram, J_k , will have different expected values. The Fisher's Kappa statistic tests whether the largest J_k can be considered different from the mean of the J_k . Critical values for the Fisher's Kappa test can be found in Fuller 1976.

The Kolmogorov-Smirnov statistic reported by PROC SPECTRA has the same asymptotic distribution as Bartlett's test (Durbin 1967). The Kolmogorov-Smirnov statistic compares the normalized cumulative periodogram with the cumulative distribution function of a uniform(0,1) random variable. The normalized cumulative periodogram, F_j , of the series is

$$F_j = \frac{\sum_{k=1}^j J_k}{\sum_{k=1}^m J_k}, j = 1, 2, \dots, m-1$$

where $m = \frac{n}{2}$ if n is even or $m = \frac{n-1}{2}$ if n is odd. The test statistic is the maximum absolute difference of the normalized cumulative periodogram and the uniform cumulative distribution function. Approximate p -values for Bartlett's Kolmogorov-Smirnov test statistics are provided with the test statistics. Small p -values cause you to reject the null-hypothesis that the series is white noise.

Transforming Frequencies

The variable FREQ in the data set created by the SPECTRA procedure ranges from 0 to π . Sometimes it is preferable to express frequencies in cycles per observation period, which is equal to $\frac{1}{2\pi}$ FREQ.

To express frequencies in cycles per unit time (for example, in cycles per year), multiply FREQ by $\frac{d}{2\pi}$, where d is the number of observations per unit of time. For example, for monthly data, if the desired time unit is years then d is 12. The period of the cycle is $\frac{2\pi}{d \times \text{FREQ}}$, which ranges from $\frac{2}{d}$ to infinity.

OUT= Data Set

The OUT= data set contains $\frac{n}{2} + 1$ observations, if n is even, or $\frac{n+1}{2}$ observations, if n is odd, where n is the number of observations in the time series or the span of data being analyzed if missing values are present in the data. See the section "Missing Values" on page 2336 for details.

The variables in the new data set are named according to the following conventions. Each variable to be analyzed is associated with an index. The first variable listed in the VAR statement is indexed as 01, the second variable as 02, and so on. Output variables are named by combining indexes with prefixes. The prefix always identifies the nature of the new variable, and the indices identify the original variables from which the statistics were obtained.

Variables that contain spectral analysis results have names that consist of a prefix, an underscore, and the index of the variable analyzed. For example, the variable S_01 contains spectral density estimates for the

first variable in the VAR statement. Variables that contain cross-spectral analysis results have names that consist of a prefix, an underscore, the index of the first variable, another underscore, and the index of the second variable. For example, the variable A_01_02 contains the amplitude of the cross-spectral density estimate for the first and second variables in the VAR statement.

Table 33.3 shows the formulas and naming conventions used for the variables in the OUT= data set. Let X be variable number *nn* in the VAR statement list and let Y be variable number *mm* in the VAR statement list. Table 33.3 shows the output variables that contain the results of the spectral and cross-spectral analysis of X and Y.

In Table 33.3 the following notation is used. Let W_j be the vector of $2p + 1$ smoothing weights given by the WEIGHTS statement, normalized to sum to $\frac{1}{4\pi}$. Note that the weights are either explicitly provided using the constant specification or are implicitly determined by the kernel specification in the WEIGHTS statement.

The subscript of W_j runs from W_{-p} to W_p , so that W_0 is the middle weight in the list. Let $\omega_k = \frac{2\pi k}{n}$, where $k = 0, 1, \dots, \text{floor}(\frac{n}{2})$.

Table 33.3 Variables Created by PROC SPECTRA

Variable	Description
FREQ	frequency in radians from 0 to π (Note: Cycles per observation is $\frac{\text{FREQ}}{2\pi}$.)
PERIOD	period or wavelength: $\frac{2\pi}{\text{FREQ}}$ (Note: PERIOD is missing for FREQ=0.)
COS_nn	cosine transform of X: $a_k^x = \frac{2}{n} \sum_{t=1}^n X_t \cos(\omega_k(t-1))$
SIN_nn	sine transform of X: $b_k^x = \frac{2}{n} \sum_{t=1}^n X_t \sin(\omega_k(t-1))$
P_nn	periodogram of X: $J_k^x = \frac{n}{2} [(a_k^x)^2 + (b_k^x)^2]$
S_nn	spectral density estimate of X: $F_k^x = \sum_{j=-p}^p W_j J_{k+j}^x$ (except across endpoints)
RP_nn_mm	real part of cross-periodogram X and Y: $\text{real}(J_k^{xy}) = \frac{n}{2} (a_k^x a_k^y + b_k^x b_k^y)$
IP_nn_mm	imaginary part of cross-periodogram of X and Y: $\text{imag}(J_k^{xy}) = \frac{n}{2} (a_k^x b_k^y - b_k^x a_k^y)$
CS_nn_mm	cospectrum estimate (real part of cross-spectrum) of X and Y: $C_k^{xy} = \sum_{j=-p}^p W_j \text{real}(J_{k+j}^{xy})$ (except across endpoints)

Table 33.3 *continued*

Variable	Variable
QS_nn_mm	quadrature spectrum estimate (imaginary part of cross-spectrum) of X and Y: $Q_k^{xy} = \sum_{j=-p}^p W_j \text{imag}(J_{k+j}^{xy}) \text{ (except across endpoints)}$
A_nn_mm	amplitude (modulus) of cross-spectrum of X and Y: $A_k^{xy} = \sqrt{(C_k^{xy})^2 + (Q_k^{xy})^2}$
K_nn_mm	coherency squared of X and Y: $K_k^{xy} = (A_k^{xy})^2 / (F_k^x F_k^y)$
PH_nn_mm	phase spectrum in radians of X and Y: $\Phi_k^{xy} = \arctan(Q_k^{xy} / C_k^{xy})$

Printed Output

By default PROC SPECTRA produces no printed output.

When the WHITETEST option is specified, the SPECTRA procedure prints the following statistics for each variable in the VAR statement:

1. the name of the variable
2. M-1, the number of two-degree-of-freedom periodogram ordinates used in the test
3. MAX(P(*)), the maximum periodogram ordinate
4. SUM(P(*)), the sum of the periodogram ordinates
5. Fisher's Kappa statistic
6. Bartlett's Kolmogorov-Smirnov test statistic
7. Approximate *p*-value for Bartlett's Kolmogorov-Smirnov test statistic

See the section “[White Noise Test](#)” on page 2339 for details.

ODS Table Names: SPECTRA procedure

PROC SPECTRA assigns a name to each table it creates. You can use these names to reference the table when you use the Output Delivery System (ODS) to select tables and create output data sets. These names are listed in the following table.

Table 33.4 ODS Tables Produced in PROC SPECTRA

ODS Table Name	Description	Option
WhiteNoiseTest	white noise test	WHITETEST
Kappa	Fishers Kappa statistic	WHITETEST
Bartlett	Bartlett's Kolmogorov-Smirnov statistic	WHITETEST

Examples: SPECTRA Procedure

Example 33.1: Spectral Analysis of Sunspot Activity

This example analyzes Wolfer's sunspot data (Anderson 1971). The following statements read and plot the data.

```

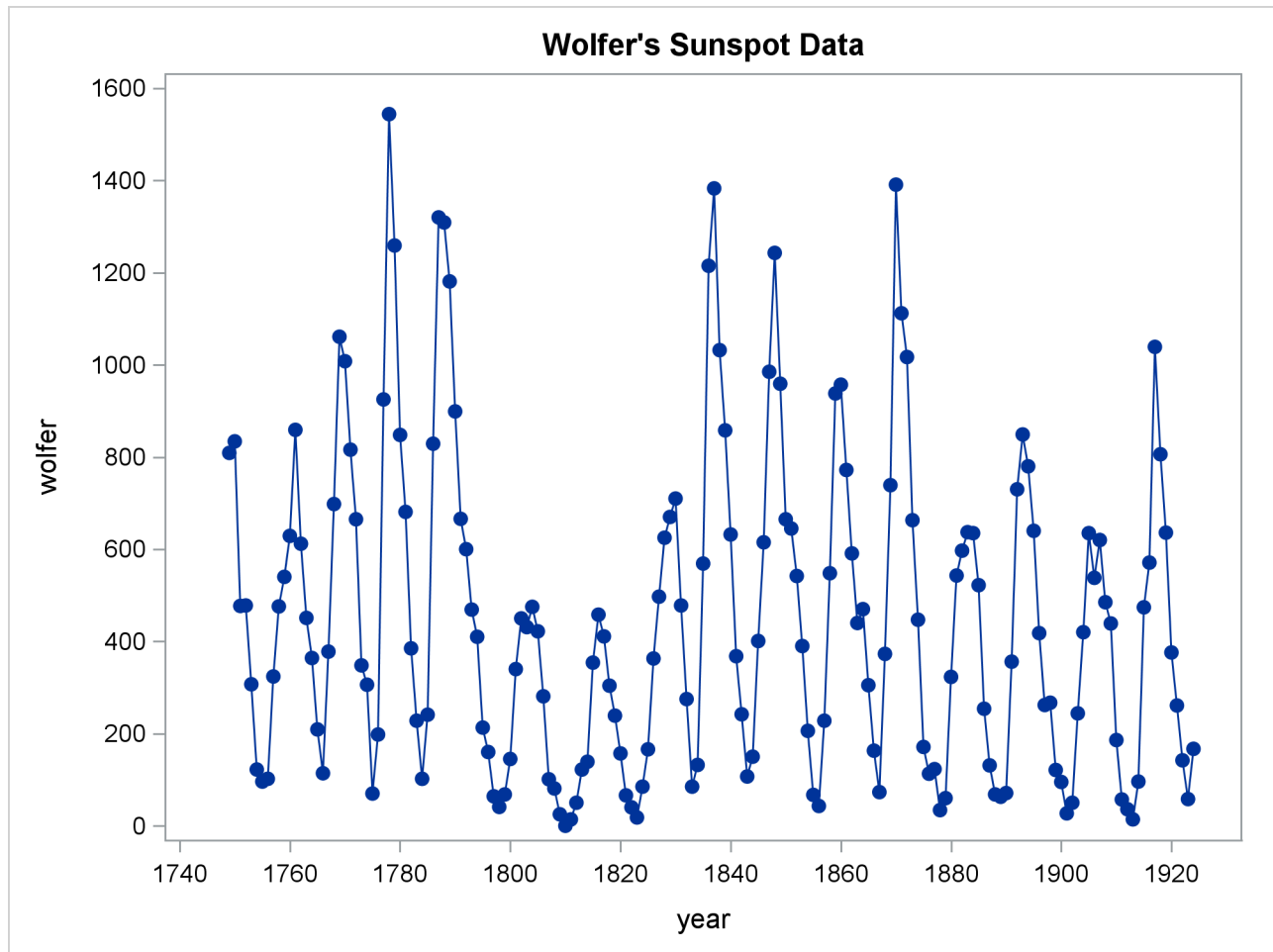
title "Wolfer's Sunspot Data";
data sunspot;
    input year wolfer @@;
datalines;
1749  809 1750  834 1751  477 1752  478 1753  307 1754  122 1755  96

... more lines ...

proc sgplot data=sunspot;
    series x=year y=wolfer / markers markerattrs=(symbol=circlefilled);
    xaxis values=(1740 to 1930 by 10);
    yaxis values=(0 to 1600 by 200);
run;

```

The plot of the sunspot series is shown in [Output 33.1.1](#).

Output 33.1.1 Plot of Original Sunspot Data

The spectral analysis of the sunspot series is performed by the following statements:

```
proc spectra data=sunspot out=b p s adjmean whitetest;
  var wolfer;
  weights 1 2 3 4 3 2 1;
run;

proc print data=b(obs=12);
run;
```

The PROC SPECTRA statement specifies the P and S options to write the periodogram and spectral density estimates to the OUT= data set B. The WEIGHTS statement specifies a triangular spectral window for smoothing the periodogram to produce the spectral density estimate. The ADJMEAN option zeros the frequency 0 value and avoids the need to exclude that observation from the plots. The WHITETEST option prints tests for white noise.

The Fisher's Kappa test statistic of 16.070 is larger than the 5% critical value of 7.2, so the null hypothesis that the sunspot series is white noise is rejected (see the table of critical values in Fuller (1976)).

The Bartlett's Kolmogorov-Smirnov statistic is 0.6501, and its approximate p -value is < 0.0001 . The small p -value associated with this test leads to the rejection of the null hypothesis that the spectrum represents white noise.

The printed output produced by PROC SPECTRA is shown in [Output 33.1.2](#). The output data set B created by PROC SPECTRA is shown in part in [Output 33.1.3](#).

Output 33.1.2 White Noise Test Results

Wolfer's Sunspot Data

The SPECTRA Procedure

Test for White Noise for Variable wolfer	
M-1	87
Max(P(*))	4062267
Sum(P(*))	21156512
Fisher's Kappa: (M-1)*Max(P(*))/Sum(P(*))	
Kappa	16.70489
Bartlett's Kolmogorov-Smirnov Statistic: Maximum absolute difference of the standardized partial sums of the periodogram and the CDF of a uniform(0,1) random variable.	
Test Statistic	0.650055
Approximate P-Value	<.0001

Output 33.1.3 First 12 Observations of the OUT= Data Set

Wolfer's Sunspot Data

Obs	FREQ	PERIOD	P_01	S_01
1	0.00000	.	0.00	59327.52
2	0.03570	176.000	3178.15	61757.98
3	0.07140	88.000	2435433.22	69528.68
4	0.10710	58.667	1077495.76	66087.57
5	0.14280	44.000	491850.36	53352.02
6	0.17850	35.200	2581.12	36678.14
7	0.21420	29.333	181163.15	20604.52
8	0.24990	25.143	283057.60	15132.81
9	0.28560	22.000	188672.97	13265.89
10	0.32130	19.556	122673.94	14953.32
11	0.35700	17.600	58532.93	16402.84
12	0.39270	16.000	213405.16	18562.13

The following statements plot the periodogram and spectral density estimate by the frequency and period.

```
proc sgplot data=b;
  series x=freq y=p_01 / markers markerattrs=(symbol=circlefilled);
run;

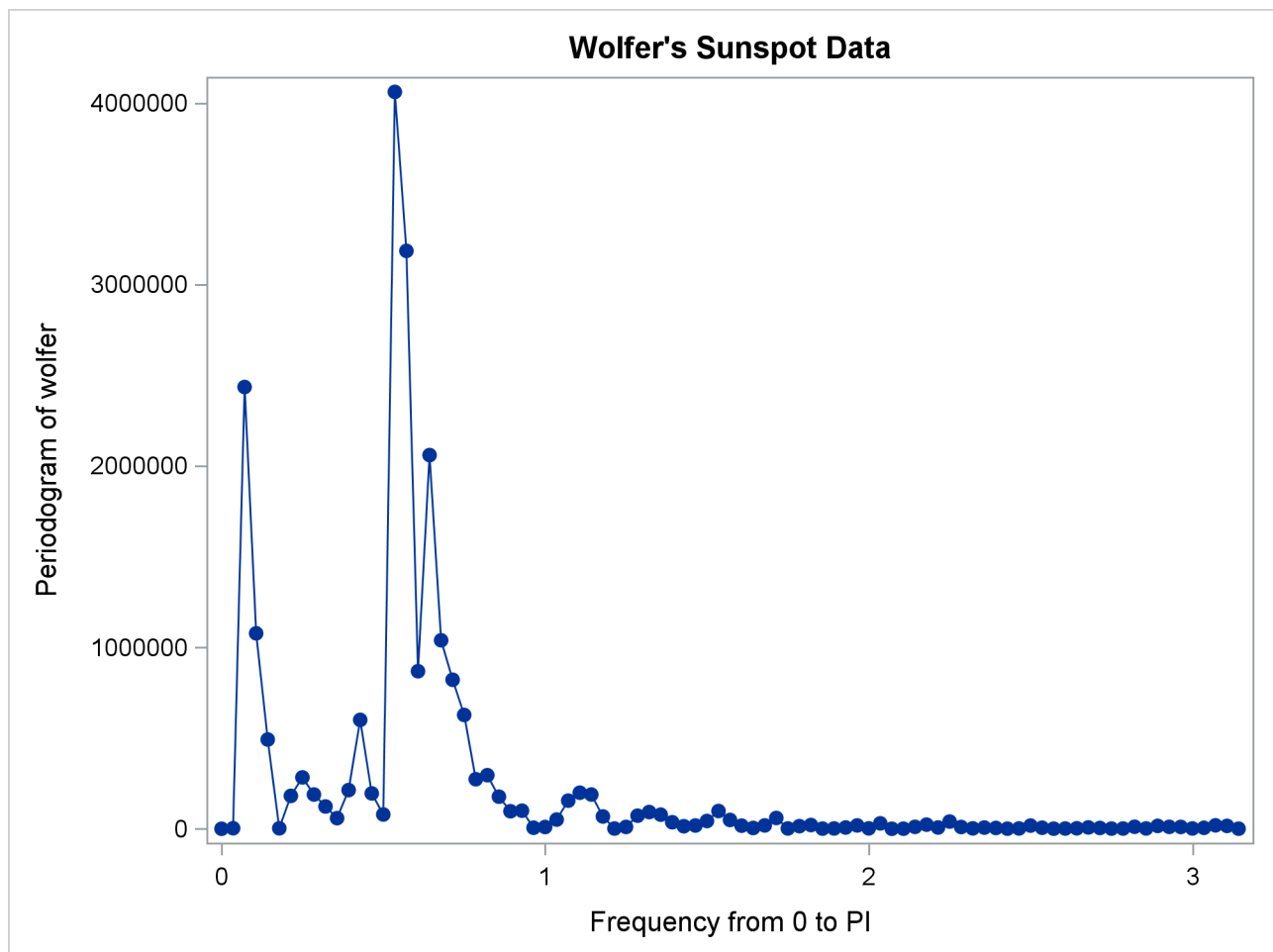
proc sgplot data=b;
  series x=period y=p_01 / markers markerattrs=(symbol=circlefilled);
run;

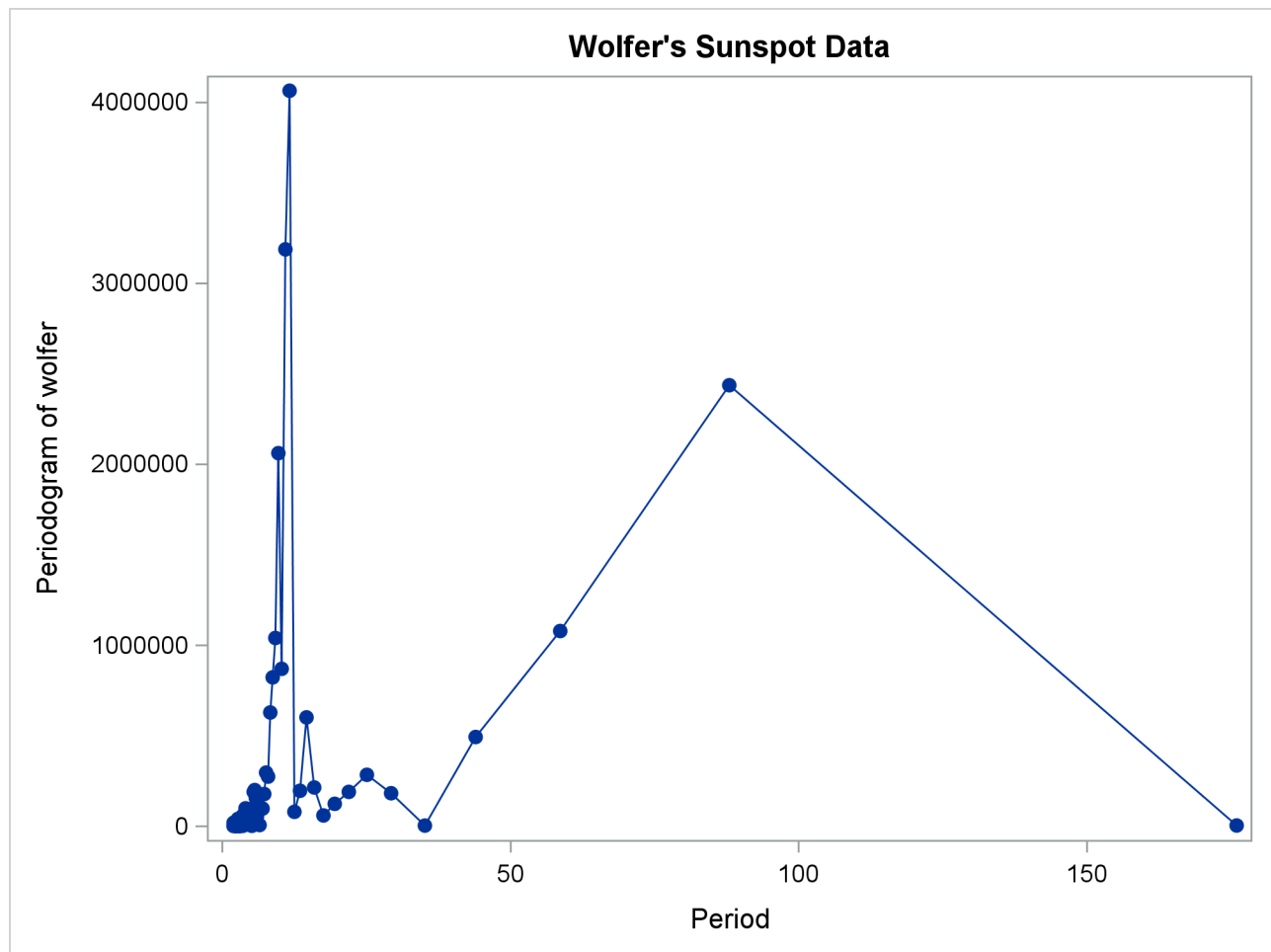
proc sgplot data=b;
  series x=freq y=s_01 / markers markerattrs=(symbol=circlefilled);
run;

proc sgplot data=b;
  series x=period y=s_01 / markers markerattrs=(symbol=circlefilled);
run;
```

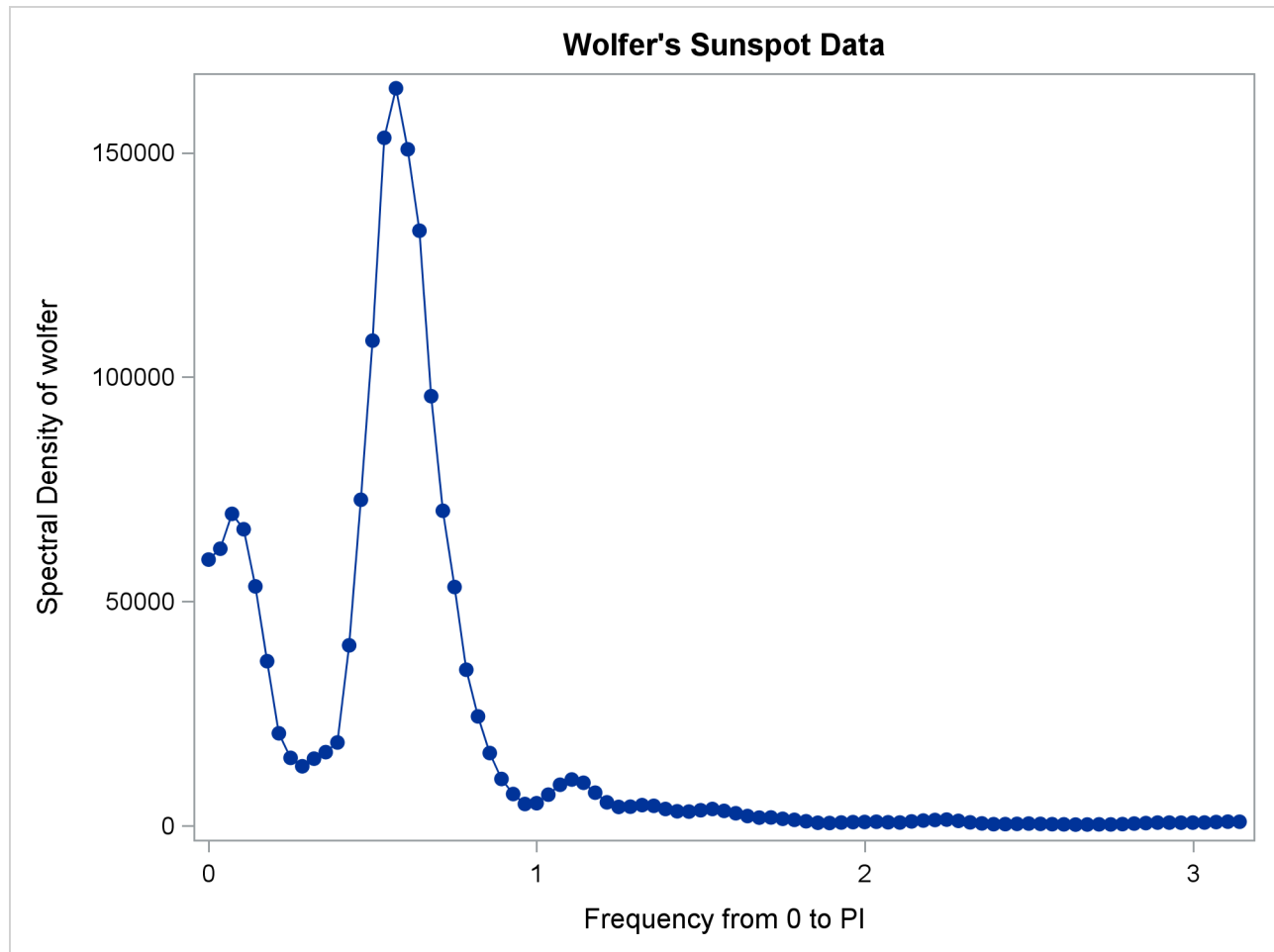
The periodogram is plotted against the frequency in [Output 33.1.4](#) and plotted against the period in [Output 33.1.5](#). The spectral density estimate is plotted against the frequency in [Output 33.1.6](#) and plotted against the period in [Output 33.1.7](#).

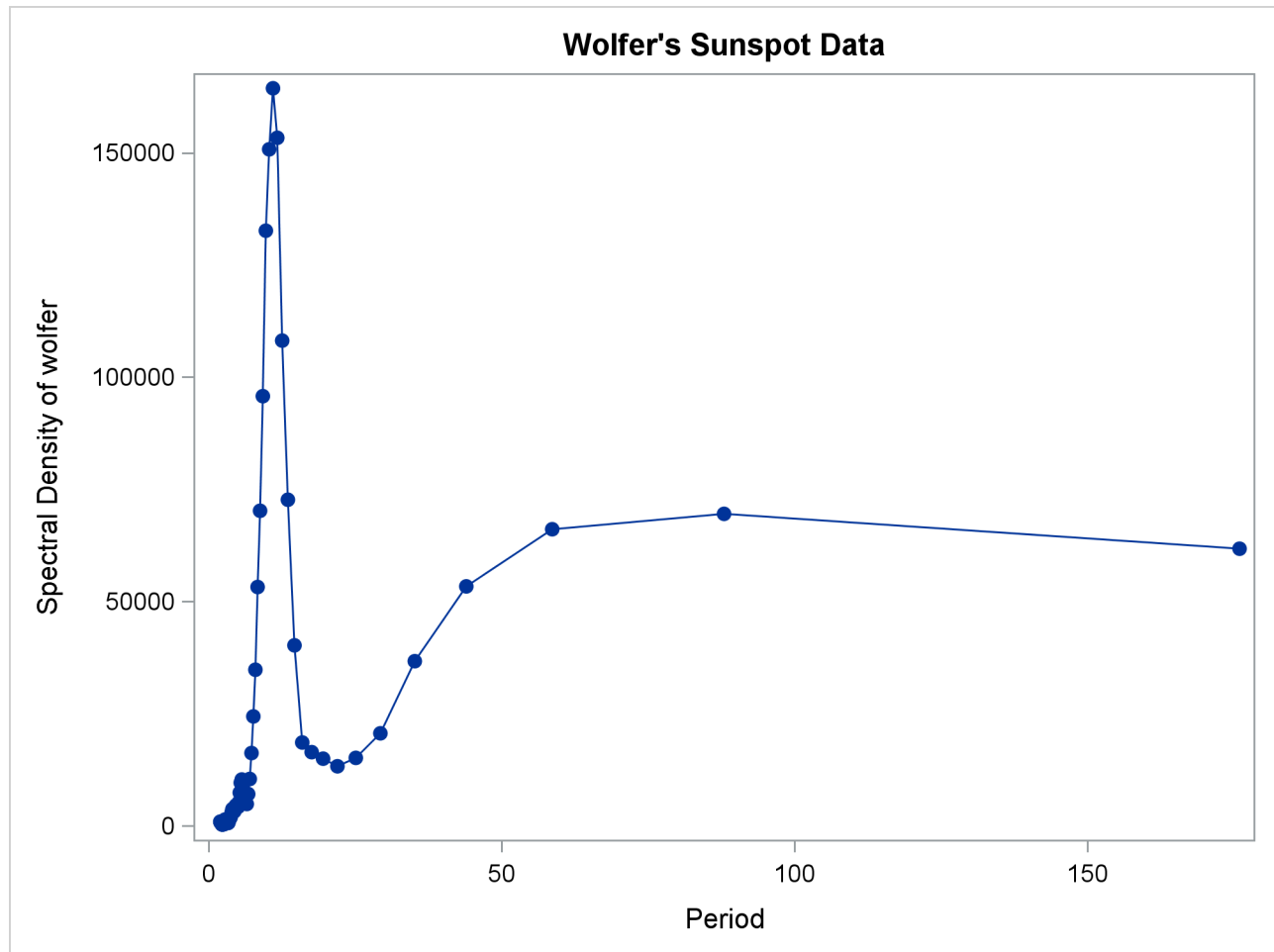
Output 33.1.4 Plot of Periodogram by Frequency



Output 33.1.5 Plot of Periodogram by Period

Output 33.1.6 Plot of Spectral Density Estimate by Frequency



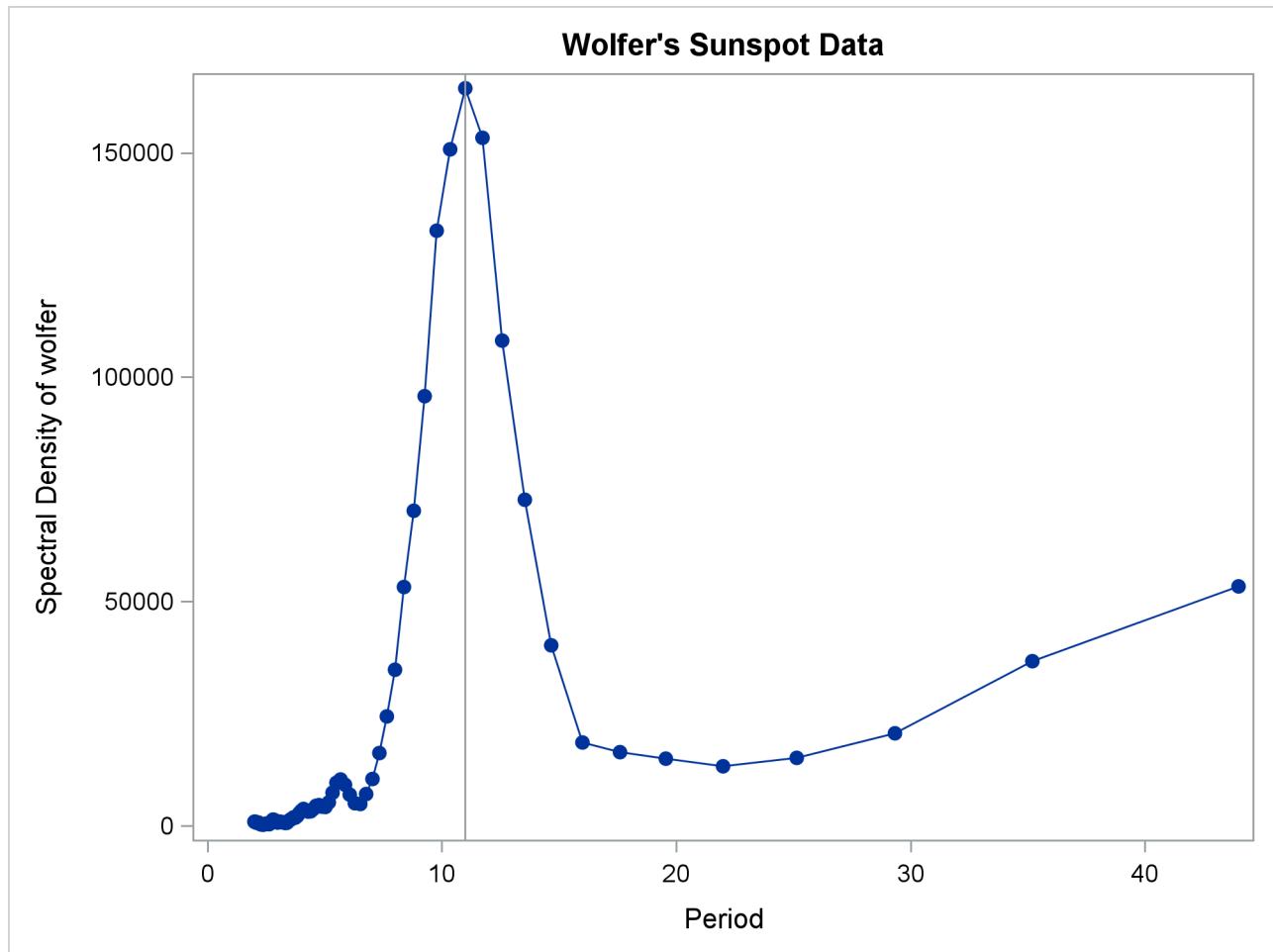
Output 33.1.7 Plot of Spectral Density Estimate by Period

Since PERIOD is the reciprocal of frequency, the plot axis for PERIOD is stretched for low frequencies and compressed at high frequencies. One way to correct for this is to use a WHERE statement to restrict the plots and exclude the low frequency components. The following statements plot the spectral density for periods less than 50.

```
proc sgplot data=b;
  where period < 50;
  series x=period y=s_01 / markers markerattrs=(symbol=circlefilled);
  refline 11 / axis=x;
run;
title;
```

The spectral analysis of the sunspot series confirms a strong 11-year cycle of sunspot activity. The plot makes this clear by drawing a reference line at the 11 year period, which highlights the position of the main peak in the spectral density.

Output 33.1.8 shows the plot. Contrast Output 33.1.8 with Output 33.1.7.

Output 33.1.8 Plot of Spectral Density Estimate by Period to 50 Years

Example 33.2: Cross-Spectral Analysis

This example uses simulated data to show cross-spectral analysis for two variables X and Y. X is generated by an AR(1) process; Y is generated as white noise plus an input from X lagged 2 periods. All output options are specified in the PROC SPECTRA statement. PROC CONTENTS shows the contents of the OUT= data set.

```
data a;
  x1 = 0; x11 = 0;
  do i = - 10 to 100;
    x = .4 * x1 + rannor(123);
    y = .5 * x11 + rannor(123);
    if i > 0 then output;
    x11 = x1; x1 = x;
  end;
run;

proc spectra data=a out=b cross coef a k p ph s;
  var x y;
  weights 1 1.5 2 4 8 9 8 4 2 1.5 1;
run;

proc contents data=b position;
run;
```

The PROC CONTENTS report for the output data set B is shown in [Output 33.2.1](#).

Output 33.2.1 Contents of PROC SPECTRA OUT= Data Set

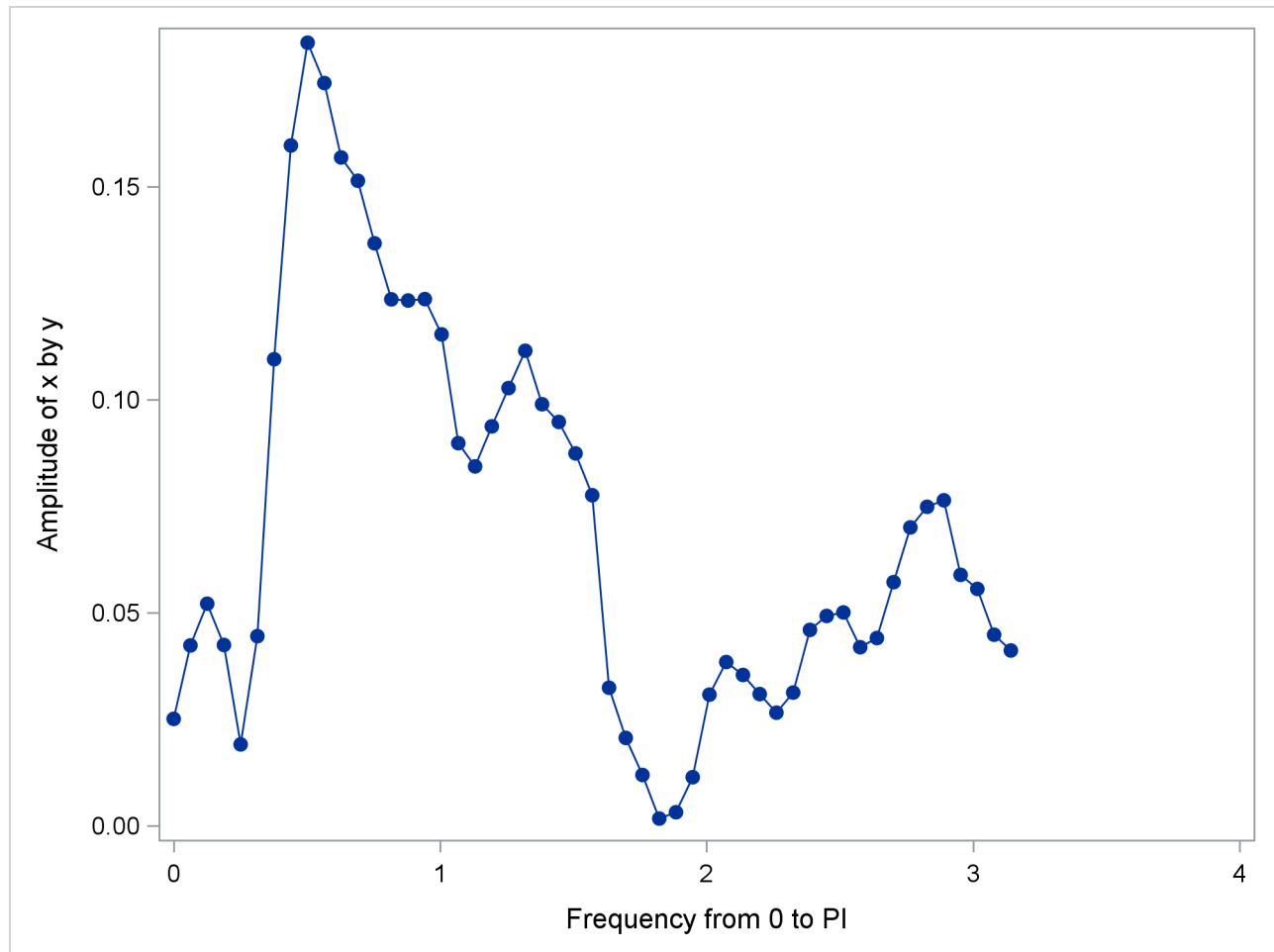
The CONTENTS Procedure

Alphabetic List of Variables and Attributes			
#	Variable	Type	Len Label
16	A_01_02	Num	8 Amplitude of x by y
3	COS_01	Num	8 Cosine Transform of x
5	COS_02	Num	8 Cosine Transform of y
13	CS_01_02	Num	8 Cospectra of x by y
1	FREQ	Num	8 Frequency from 0 to PI
12	IP_01_02	Num	8 Imag Periodogram of x by y
15	K_01_02	Num	8 Coherency**2 of x by y
2	PERIOD	Num	8 Period
17	PH_01_02	Num	8 Phase of x by y
7	P_01	Num	8 Periodogram of x
8	P_02	Num	8 Periodogram of y
14	QS_01_02	Num	8 Quadrature of x by y
11	RP_01_02	Num	8 Real Periodogram of x by y
4	SIN_01	Num	8 Sine Transform of x
6	SIN_02	Num	8 Sine Transform of y
9	S_01	Num	8 Spectral Density of x
10	S_02	Num	8 Spectral Density of y

The following statements plot the amplitude of the cross-spectrum estimate against frequency and against period for periods less than 25.

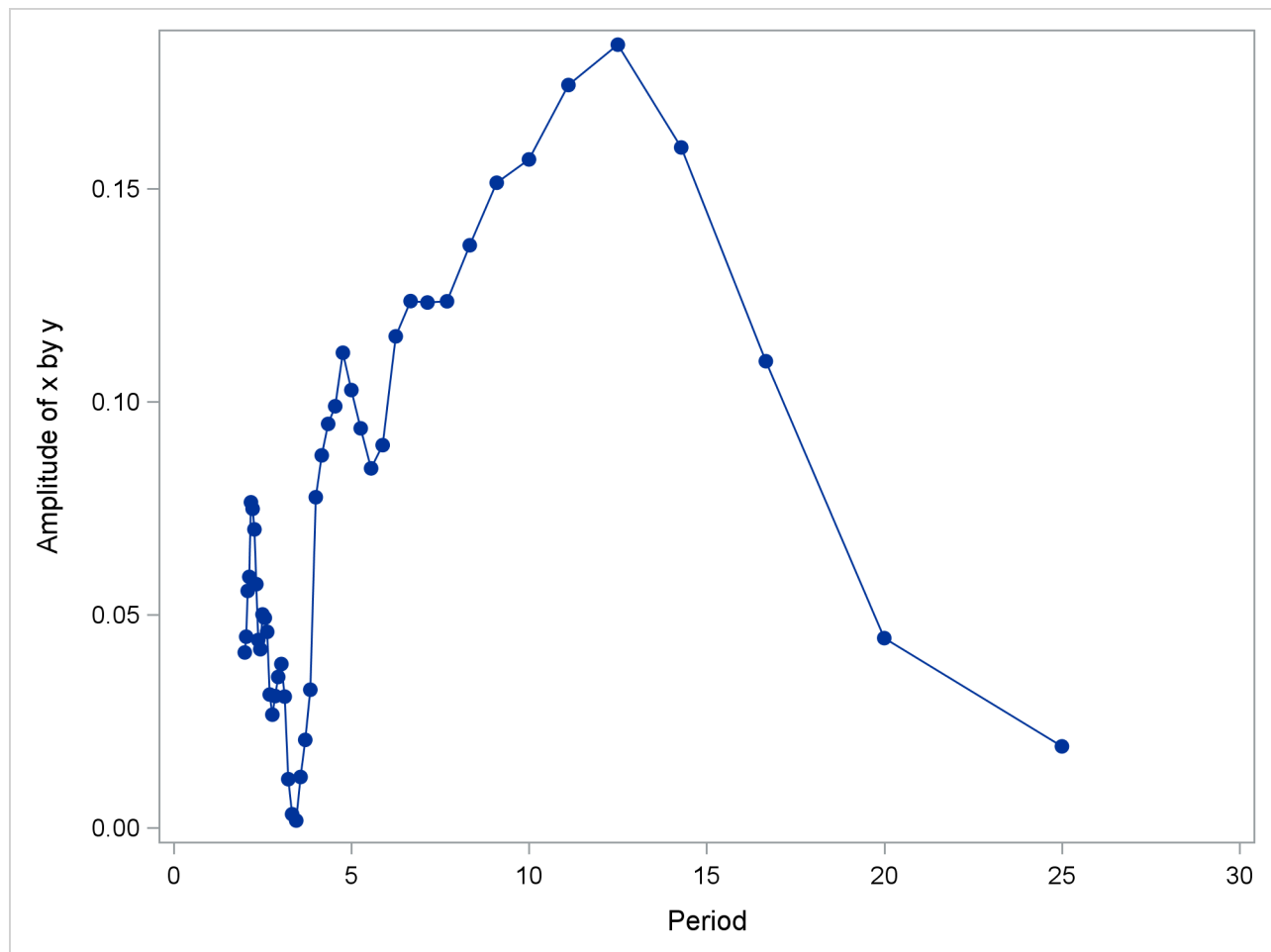
```
proc sgplot data=b;
  series x=freq y=a_01_02 / markers markerattrs=(symbol=circlefilled);
  xaxis values=(0 to 4 by 1);
run;
```

The plot of the amplitude of the cross-spectrum estimate against frequency is shown in [Output 33.2.2](#).

Output 33.2.2 Plot of Cross-Spectrum Amplitude by Frequency

The plot of the cross-spectrum amplitude against period for periods less than 25 observations is shown in [Output 33.2.3](#).

```
proc sgplot data=b;
  where period < 25;
  series x=period y=a_01_02 / markers markerattrs=(symbol=circlefilled);
  xaxis values=(0 to 30 by 5);
run;
```

Output 33.2.3 Plot of Cross-Spectrum Amplitude by Period

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